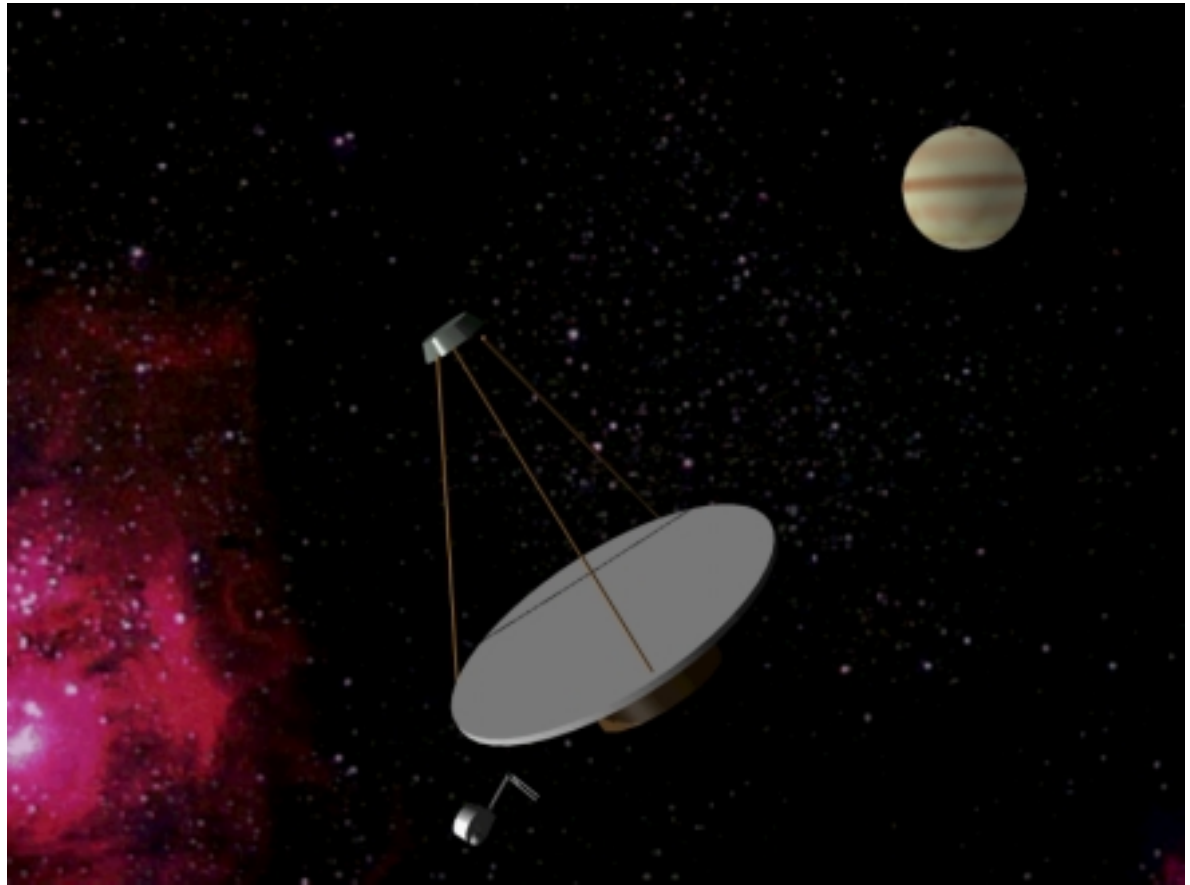


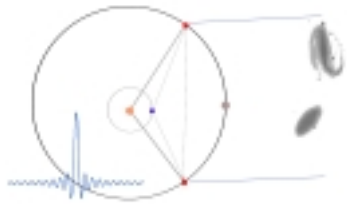
PLANETARY-SCALE ASTRONOMICAL BENCH

**NIAC Phase I Study
November 1999**



Timothy L. Howard, *Principal Investigator*
Lee Gutheinz, Jack Sanders-Reed, Carl Tuttle, Bill Witt

SVS, Inc.



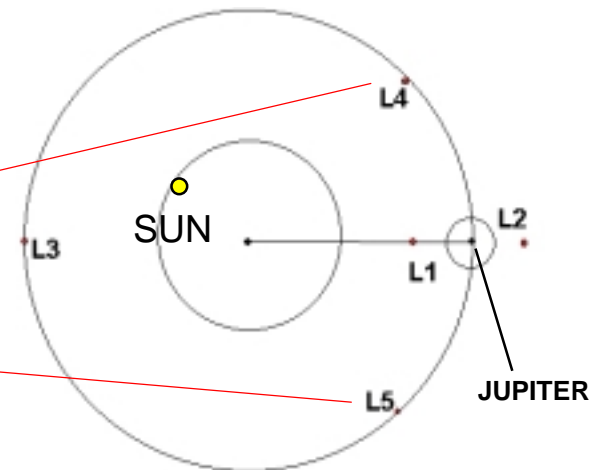
PAB: WHAT IS IT ?

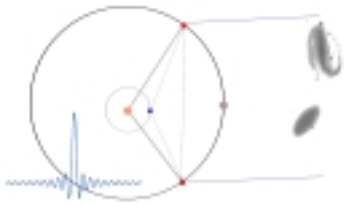
AN ASTRONOMICAL OBSERVATORY

- USING JOVIAN LAGRANGE POINTS AS LONG-TERM SITES FOR ASTRONOMICAL INSTRUMENTS
- LOCATIONS OF NATURAL ORBITAL STABILITY
- USE SOLAR-SYSTEM SCALE BASELINES FOR COOPERATIVE MEASUREMENTS
- CONCEPT STUDY: EXAMINE POTENTIAL APPLICATIONS, BASIC FEASIBILITY ISSUES, SYSTEM CONCEPTS AND TECHNOLOGY ISSUES
- ENVISION NOMINAL TIMEFRAME FOR DEVELOPMENT
10-40 YEARS

PAB “STATIONS”

Nomenclature:
JL4, JL5 (generically, JLx)





PAB BASIC INFO

**STATIONS -- JL4 (leading), JL5 (trailing)
60 deg. ahead/behind JUPITER
E1 (Earth locus)**

**ORBIT: 5.2 AU (radius) PERIOD 11.86 yr
[1 AU = 778 million km]**

JL4-JL5 DISTANCE = 9.01 AU (nominal)

JLx-E1 DISTANCE = 4.2 - 6.2 AU (periodic)

SOLAR IRRADIANCE AT 5.2 AU ~ 50 W/m²

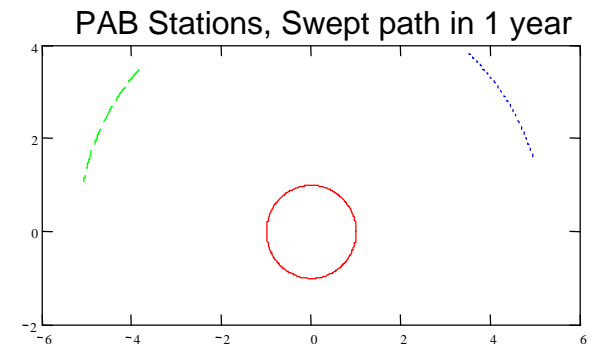
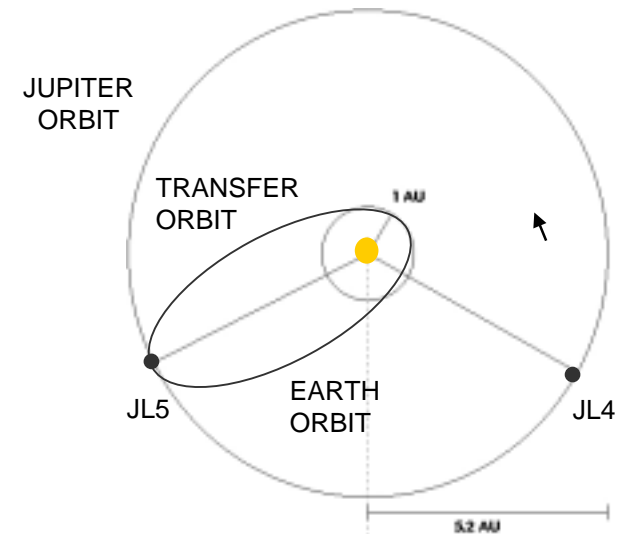
LIGHT-TRAVEL TIME:

JLx-E1 : 34.9 - 51.6 min.

JL4-JL5 : 74.9 min.

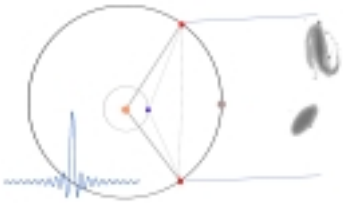
**HOHMANN ORBIT
TRANSFER TIME = 2.6 yr (one-way)**

LAUNCH WINDOW ~ 13 MONTHS



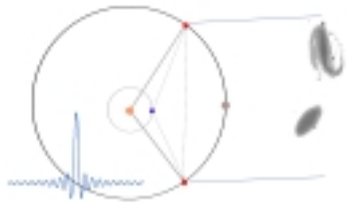
(looking down from North Ecliptic Pole; rotation CCW)

PAB APPLICATIONS

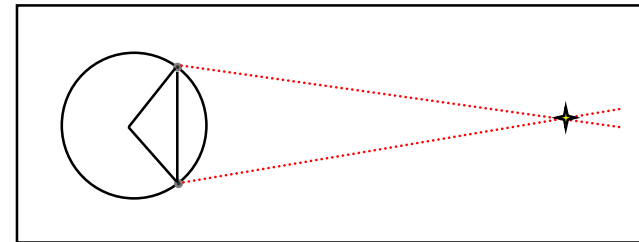


PLANETARY-SCALE ASTRONOMICAL BENCH

- **PARALLAX-BASED ASTROMETRY**
- **CONVENTIONAL ASTRONOMICAL IMAGING & SPECTROSCOPY**
- **MICROLENSING**
- **LONG-BASELINE INTERFEROMETRY
RADIO, AND OPTICAL/IR
FOR ASTROMETRY AND SYNTHESIS IMAGING**
- **SOLAR SYSTEM SCIENCE
VIA LOCAL PROBES TO TROJAN ASTEROID POPULATION
SENSORS TO MONITOR INTERPLANETARY ENVIRONMENT**
- **GRAVITATIONAL WAVES**
- **OTHER LONG-BASELINE DYNAMICS STUDIES**

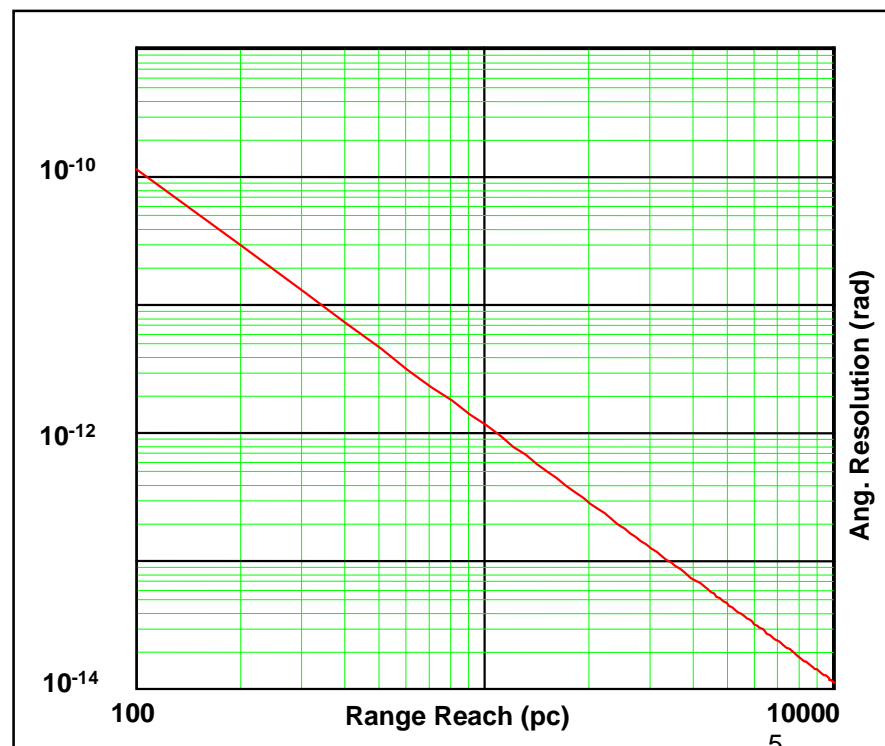


PAB SCIENCE: ASTROMETRY



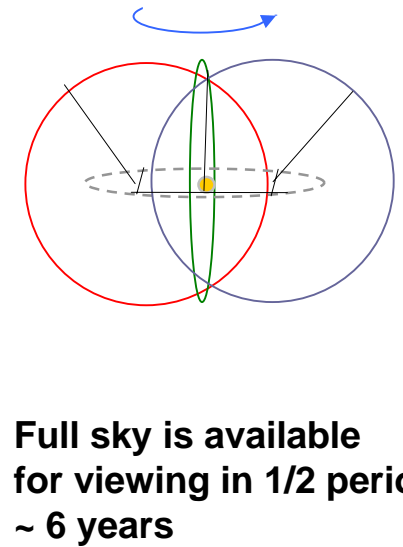
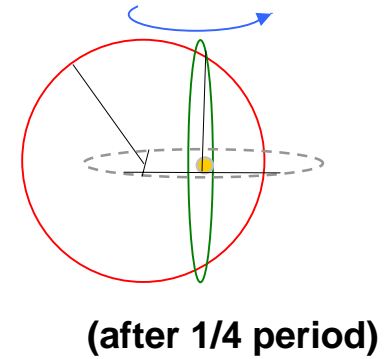
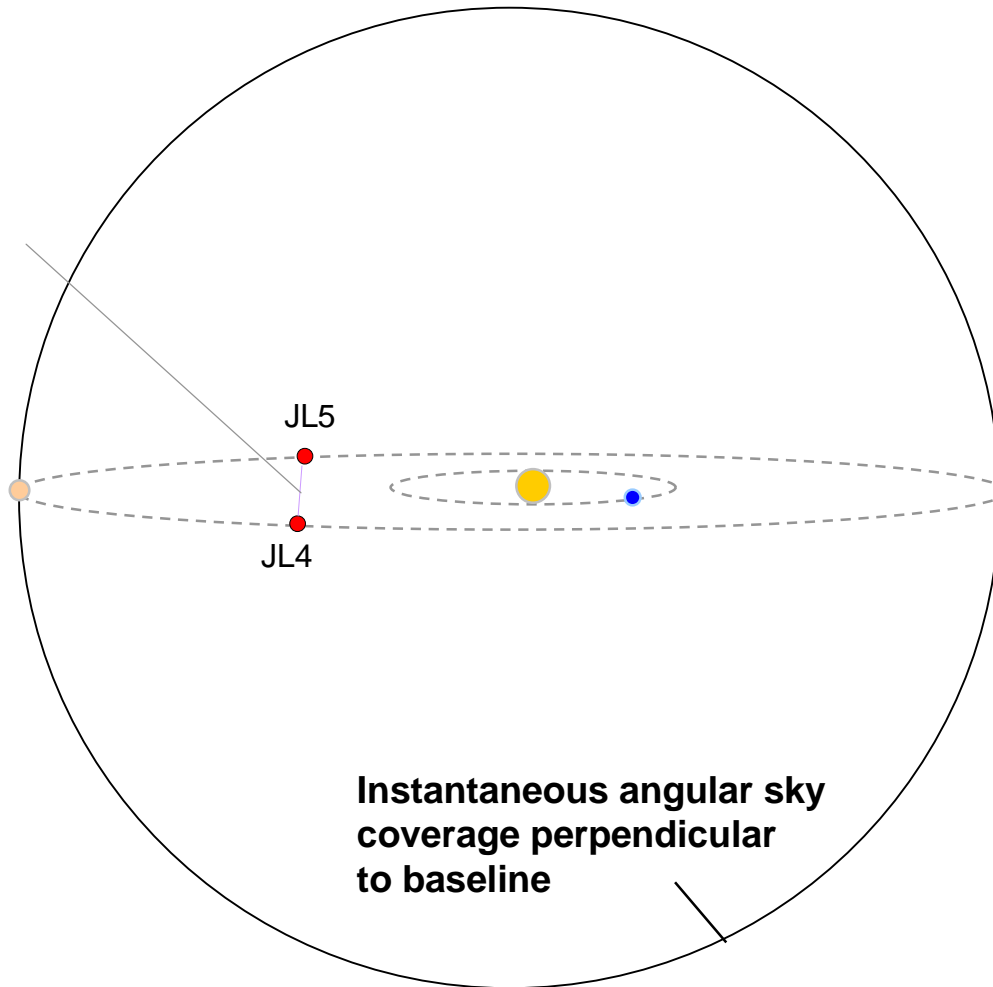
PLANETARY-SCALE ASTRONOMICAL BENCH

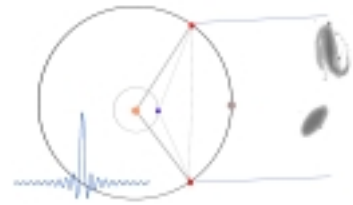
- **~9AU FOR PARALLAX-BASED RANGE**
- **RANGE REACH DEPENDS ON ANGULAR RESOLUTION**
 - AT 10^{-12} RAD (0.2 MICROARCSEC), $R > 10$ MPC
(c.f. ~8500 PC TO GAL. CENTER, 600000 PC TO M31)
- **CURRENT TRENDS**
 - "FEW" MICROARCSEC
~ 2005 (SIM, GAIA)
- **ASTROMETRIC 'DEPTH SENSING' ENABLED VIA SIMULTANEOUS IMAGING**
 - RANGE RESOLUTION
0.1 PC at ~1000 PC



SKY COVERAGE

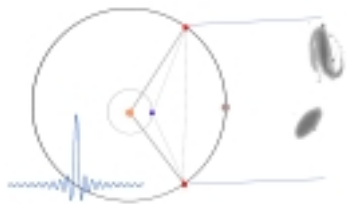
PLANETARY-SCALE ASTRONOMICAL BENCH





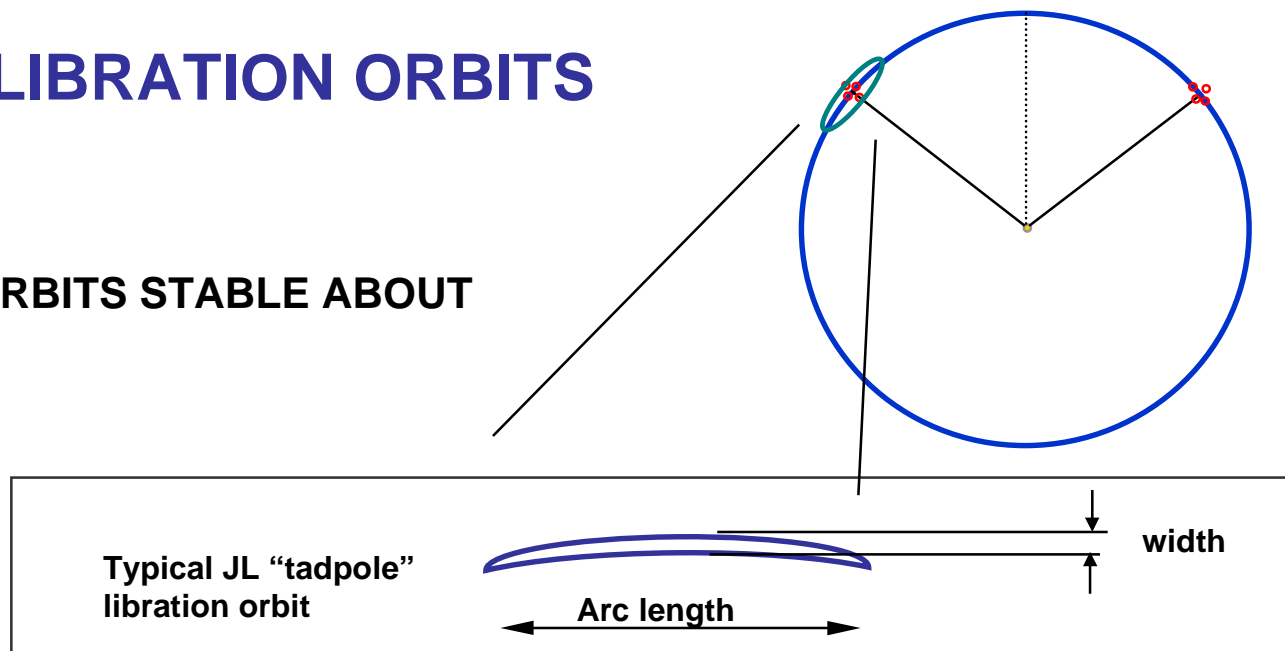
APERTURE SYNTHESIS

- USE OF JLx POINTS ANALYZED FOR APERTURE SYNTHESIS
- INTERFEROMETRY ACROSS FULL 9 AU BASELINE THEORETICALLY POSSIBLE BUT IS SEVERELY SIGNAL-STARVED
 - LIMITS CAN BE DERIVED IN TERMS OF AN “AREA-TIME” PRODUCT (COLLECTING AREA x INTEGRATION PERIOD)
 - SCALES AS $(\text{BASELINE})^2$ AND DEPENDENT ON SOURCE TEMPERATURE AND WAVELENGTH
- TYPICAL VALUE: T(blackbody) @ 6000K (sunlike star)
BASELINE 10 AU : $A\Delta T = 10^{15} \text{ m}^2\text{-sec}$, 1 micron
- PRACTICAL INTEGRATION TIMES LIMITED TO $\sim 10^5 \text{ sec}$
 - ROTATION OF BASELINE CHANGES RESOLUTION SIZE
- EFFECTIVE USE OF LAGRANGE POINTS
 - > “LIBRATION-PAIR” ORBITS



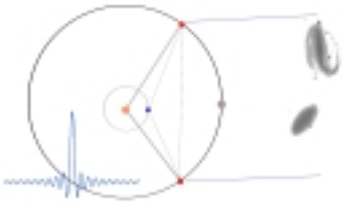
LIBRATION ORBITS

- LIBRATION ORBITS STABLE ABOUT JLx POINTS



- JLx WIDTHS UP TO ~ 0.16 AU, ARC LENGTHS ~ 60 DEGREES
- LIBRATION PERIOD ~ 150 YEARS, MOTION ABOUT JLx IS SIMPLE HARMONIC TO FIRST ORDER IN RADIUS AND LONGITUDE
- PLACE PAIRS (OR LARGER MULTIPLES) OF APERTURES IN SYMMETRIC POSITIONS ABOUT ONE OR BOTH JLx POINTS
- NATURAL ORBITAL MOTION PLUS SLOW LIBRATION DRIFT DEVELOPS (u,v) COVERAGE OVER TIME

SNR AND RADIOMETRY FOR INTERFEROMETRY



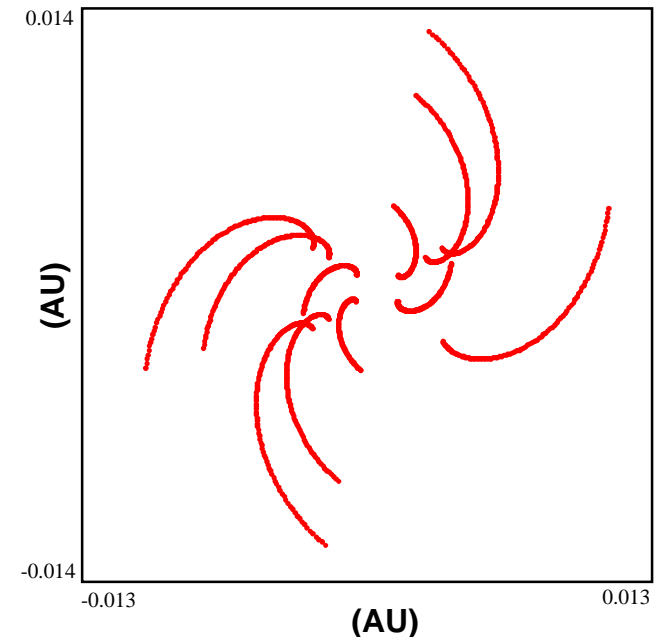
- CALCULATE BOTH A RADIOMETRIC AND A RESOLUTION RANGE REACH
 - RADIOMETRIC RANGE REACH SET BY SOURCE FLUX, SPECTRAL BANDPASS CONSISTENT WITH APERTURE SYNTHESIS
 - RESOLUTION RANGE REACH (APERTURE SYNTHESIS) SET BY BASELINE TO JUST RESOLVE AT WAVELENGTH
 - USE 1 PHOTON/SEC AS LIMITING CASE, SELECT SEPARATION AND/OR APERTURE TO MATCH THEM
- RESULTS SUGGEST 0.01 AU, 100M APERTURE A CLOSE MATCH (2X)
 - $A\Delta t \sim 2 \times 10^8$ AT 1 micron
 - ANGULAR RESOLUTION 10^{-15} rad AT 1 micron
 - EXCEEDS ANY CURRENT PLANS
 - RESOLVE SUNLIKE STARS AT ~ 10 Mpc

BASELINE COVERAGE

- LIBRATION-PAIR MOTION DEVELOPS REASONABLE BASELINE COVERAGE
 - USE OF ITERATIVE DECONVOLUTION ALGORITHMS TO CLEAN THE IMAGE
-
- LIBRATION ORBITS ABOUT OTHER SOLAR SYSTEM LAGRANGE POINTS
 - JLx POINTS HAVE LARGER REGIONS OF STABILITY, LOWER ANGULAR RATES

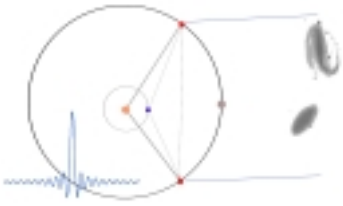
-- .002 AU Max (Earth)
-- < .001 AU (Mars)

RATES 6x - 8x LOWER
FOR JLx vs ELx, MLx

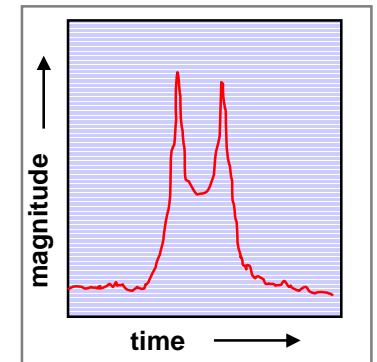
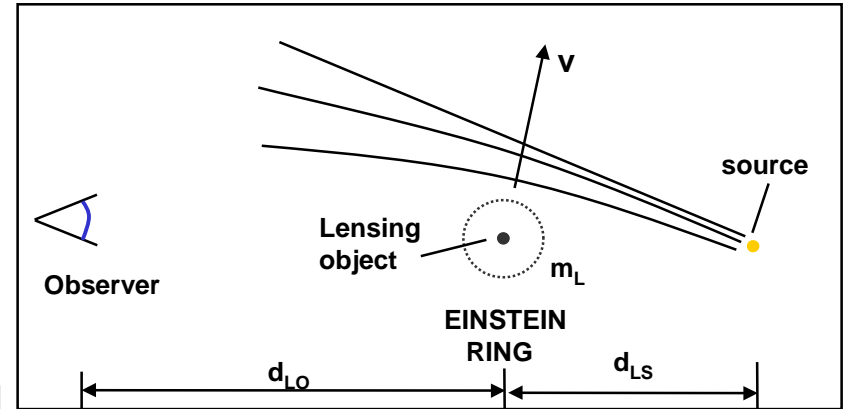


- 0.001 AU initial spacing
- 2 ea. (JL4A+B, JL5A+B)
- short λ scan $0.75 \cdot \lambda_0$
- long λ scan $1.35 \cdot \lambda_0$
- 1/3 ORBIT

PAB SCIENCE : MICROLENSING



- **GRAVITATIONAL LENSING WHERE SOURCE IS NOT RESOLVED; CAUSES BRIGHTNESS CHANGE**
- **CURRENTLY GROUND-BASED WITH ~ 1- 2 M APERTURES, MULTI-BAND PHOTOMETRY AT ~1-2% PRECISION**
- **TIME SIGNATURE OF LENSING OBJECT IS SYMMETRIC AND ACHROMATIC**
- **TIMESCALES: ~DAYS (UP TO 1 MONTH) FOR STELLAR-MASS OBJECTS; HOURS TO ~ 1 DAY FOR PLANET OBJECTS**
- **SINGLE OBSERVATIONS DO NOT RESOLVE LENSING OBJECT MASS, DISTANCE, MOTION INDEPENDENTLY**
- **PARALLAX OBSERVATIONS + DOPPLER SPECTRAL ANALYSIS COULD PROVIDE FULL SOLUTION**

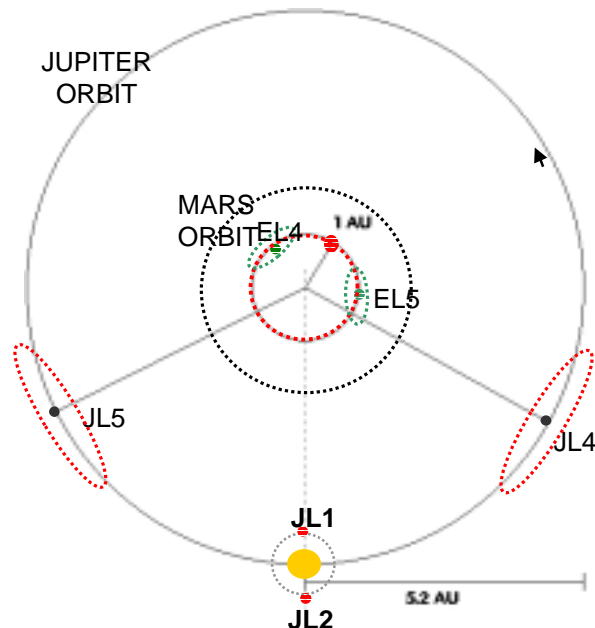


SYSTEM CONCEPTS & SIZING (1)

PLANETARY-SCALE ASTRONOMICAL BENCH

- **STATION LOCATIONS: JL4, JL5, EL4**
- **MULTIPLE INSTRUMENT POSSIBILITIES ON-STATION**
- **SYSTEM BASING OPTIONS:**

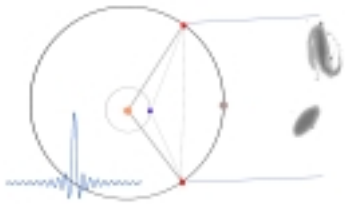
- FREE-FLYERS] - HYBRID
- MONOLITHIC
- ASTEROID (?)



CANDIDATE INSTRUMENTS

APPLICATION	**	SIZE	JL4	JL5
CONVENTIONAL ASTRONOMY		10m	X	
ASTROMETRY	x	>2m	X	X
APERTURE SYNTHESIS	x	>10m	XX	XX
MICROLENSING	x	>1m	X	X
ASTEROID STUDIES		>1m	X	
ROBOTICS (EXPLORATION)	x	n/a	X	X

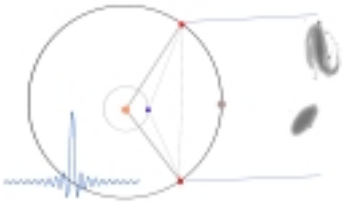
** = dedicated aperture



SYSTEM CONCEPTS & SIZING (1)

- **TOTAL SYSTEM MASS USING FULL “CANDIDATE INSTRUMENTS” LIST**
 - **ROM ESTIMATE FOR A “REALITY CHECK”**
~30000 kg ea. STA. (JL4, JL5)
 - **SIZING RULES: AREAL DENSITY 10 kg/m² (NGST),
+ INSTRUMENT ELECT./METROLOGY (100/200 kg)
+ SPACECRAFT OVERHEAD (3.3x P/L SUPPORT + 0.55xTOTAL)**
 - **INCL. STATIONKEEPING FUEL ALLOC. BUT NOT TRANSFER STAGE
CONSERVATIVE, CAN PROBABLY REDUCE BY $\geq 2X$**
- **LOGISTICS ADVANTAGES TO SOME SHARED SERVICES / FUNCTIONS**
 - **CENTRAL POWER AND DATA RELAY, STATION-STATION METROLOGY**
 - **LOCAL POWER, COMM AND FINE-SCALE METROLOGY**
- **SYSTEM WOULD BE DEVELOPED / BUILT UP INCREMENTALLY**

ENVIRONMENT / DISTURBANCES



SMALL-SCALE DISTURBANCES:

-- METEOROID FLUX	$\sim 10^{-5}$ N
-- RESIDUAL GAS	10^{-10}
-- SOLAR WIND	10^{-6}
-- RADIATION PRESSURE	10^{-5}
-- COSMIC RAYS	10^{-9}

GRAVITATIONAL PERTURBATIONS:

-- SATURN,	10^{-7} m/s ²
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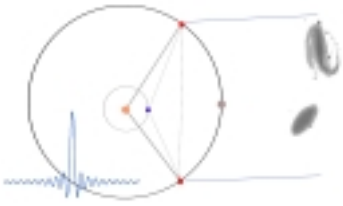
SMALL-SCALE METROLOGICAL EFFECTS:

-- STELLAR ABERRATION	~ 40 μ rad
-- GRAVITATIONAL BENDING OF LIGHT	10 prad - 0.1 μ rad

DELTA-V REQUIREMENTS

-- STATIONKEEPING (~ 1 /DAY)	< 0.01 m/s/day
-- ORBIT ADJUSTMENT (~ 1 /YR) (LIBRATION TRUNCATED)	~ 15 m/sec/yr

TECHNOLOGY DRIVERS (1)



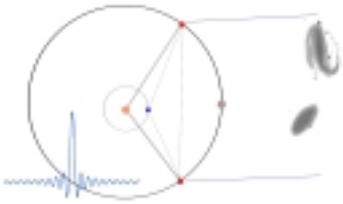
LARGE OPTICS

- PAB SHOULD HAVE $> 10\text{M}$ OPTICS, 100M DESIRABLE
- $10 - 100\text{ M}$ SPACE OPTICS DEVELOPMENT, SEVERAL GROUPS
- RECENT LARGE OPTICS DESIGNS: FREE-FLYER STRUCTURES
- AREAL DENSITY: WORKING ON 15 kg/m^2 ;
EXPECT 10 BY ~ 2005 ;
WOULD LIKE < 1 FOR PAB

POSITION & TIME MEASUREMENT

- $\sim 1\text{m}$ STATION-STATION TO SUPPORT PARALLAX TO $> 10000\text{pc}$
- SUB-WAVELENGTH POSITION MEASUREMENT FOR APERTURE SYNTHESIS
- TIME RESOLUTION TO 10^{-15} SEC ; LIKELY EXTRAPOLATION OF CURRENT TECHNOLOGY (10X)

TECHNOLOGY DRIVERS (2)



SEPARATED-SPACECRAFT INTERFEROMETRY

- NOT CURRENTLY ACHIEVABLE AT $>\sim 100\text{m}$ BASELINES
- PREVIOUSLY PROPOSED MISSIONS MAY EXPAND WORKING BASELINES TO FROM $\sim 1\text{km}$ TO $\sim 1000\text{ km}$ (ST3, DARWIN, PI)
- VERY-LONG-BASELINE OPTICAL/IR INTERFEROMETRY NEEDS DEVELOPMENT OF L.O./CORRELATION METHODS (c.f. RADIO VLBI)
- ALTERNATIVE WOULD BE DIRECT AMPLITUDE/PHASE RECORDING

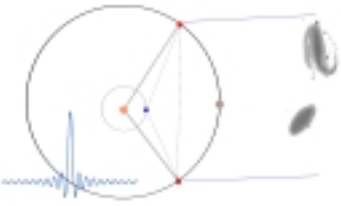
SPACECRAFT ROBOTICS

- PAB CAN TAKE ADVANTAGE OF LIKELY DEVELOPMENTS IN “SMARTER” SPACECRAFT
- AUTONOMOUS NAV, PROXIMITY OPS/DOCKING, INSTRUMENT SERVICING, DIAGNOSIS/SELF-TEST

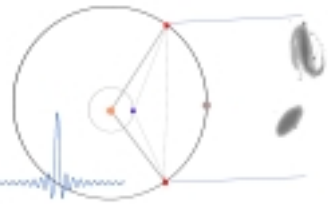
OTHER SPACECRAFT SYSTEMS

- FINE POINTING/TRACKING, OPTICAL COMM/COMPUTING
- MORE EFFICIENT PROPULSION

EXOSOLAR PLANET DETECTION



- **NARROW-ANGLE ASTROMETRY**
 - MEASURE REFLEX MOTION OF STAR ABOUT BARYCENTER
 - $3 \mu\text{arcsec}$ FOR $1 M_{\text{EARTH}}, 1 M_{\text{SUN}}, 1\text{AU}$ @ 1 pc , SCALES ~ RANGE
- **MICROLENSING**
 - PARALLAX EFFECT --> TIME OFFSET BETWEEN LENSING SIGNATURE, ~ 10s OF DAYS
 - SIZE OF EFFECT SCALES WITH PLANET, EARTH-SIZED PLANETS DETECTABLE
 - USE IN COORDINATION WITH ASTROMETRY
- **GOOD LOCATION FOR A NULLING INTERFEROMETER (OUTSIDE ZODIACAL CLOUD)**
- **OTHER METHODS POSSIBLE, e.g., OCCULTATION**



ASTEROID STUDIES & PRECURSOR MISSIONS

CATALOGUED ASTEROIDS (C. 12/98)
(VIEW FROM ABOVE ECLIPTIC)

Trojan groups

Main belt

Jupiter

Viewpoint: 3 billion km
(20 AU) above the sun



- TROJANS REPRESENT EARLY SOLAR SYSTEM MATERIAL
- LIKELY ANALOGS: “DEAD COMETS”; CARBONACEOUS CHONDRITES, POSSIBLE RESIDUAL WATER ICE
- SEVERAL HUNDRED KNOWN, SEVERAL INTERACTING GROUPS; TYPICAL SIZE ~ 15 km
- CONSIDER SINGLE LAUNCH TO ONE JL POINT (JL4 STRAWMAN)
- ROBOTIC EXPLORATION OF TROJAN ASTEROIDS & LOCAL ENVIRONMENT
- SHOULD INCLUDE ASTRONOMICAL INSTRUMENTATION